

$$gmst = 6.6974243242 + 0.0657098283 * (n) + hour \dots\dots\dots (2. 11)$$

$$lmst = (gmst * 15 + long) * (\pi/180) \dots\dots\dots (2. 12)$$

$$\omega = lmst - ra \dots\dots\dots (2. 13)$$

$$\theta z = \cos^{-1}[\cos \Phi \cos \omega \cos \delta + \sin \delta \sin \Phi] \dots\dots\dots (2. 14)$$

$$\gamma = \tan^{-1} \left(\frac{-\sin \omega}{\tan \delta \cos \Phi - \sin \Phi \cos \omega} \right) \dots\dots\dots (2. 15)$$

$$Parallax = \frac{EarthMeanRadius}{AstronomicalUnit} * \sin \theta z \dots\dots\dots (2. 16)$$

$$\theta_t = \theta_z + Parallax \dots\dots\dots (2. 17)$$

$$\rho = \tan^{-1} \frac{\sin(\gamma)}{\tan(180/\pi) - \theta_t} \dots\dots\dots (2. 18)$$

Where:

jd: Julian Day

m: Month {required from the PLC}

y: Year {required from the PLC}

d: Day {required from the PLC}

hour: Hour in Universal Time {required from the PLC}

n: The difference between current Julian Day and Julian Day
2000)

L: Mean longitude of the Sun

g: Mean anomaly of the Sun

l: Ecliptic longitude of the Sun